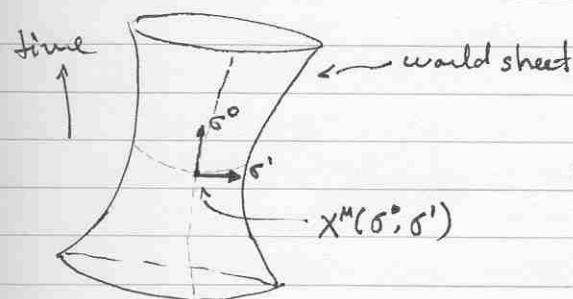


Relativistic String

Particle:

$$s = m \int d\tau = m \int \sqrt{\dot{X}^\mu \dot{X}_\mu} d\lambda \rightarrow S[e, X] = \frac{m}{2} \int [e^{-1} \dot{X}^\mu \dot{X}^\nu g_{\mu\nu} + e] d\lambda$$

omit e for massless case



$$S = T \int d^2 A : \text{Nambu - Goto action (nothing more than the worldsheet area)}$$

string tension

index: $a, b, \dots = 1, 2$

$\mu, \nu, \dots = 1, 2, 3, 4$

$\partial_\alpha X^\mu \rightarrow$ tangent vector to the worldsheet

in analogy to the particle, we ~~see~~ ^{rel}:

$$S[X^\mu, h^{ab}] = \frac{T}{2} \int \sqrt{h} h^{ab} \underbrace{\partial_\alpha X^\mu \partial_\beta X^\nu g_{\mu\nu}(X)}_{\text{call it } H_{ab}} d^2 \sigma \leftarrow \text{Polyakov action}$$

(non-linear σ model)

if replace $h_{ab} \rightarrow \Omega^2 h_{ab}$ (a Weyl or conformal transformation)

$$\text{so: } \det h \rightarrow \Omega^4 \det h, \quad \sqrt{h} \rightarrow \Omega^2 \sqrt{h}, \quad \left\{ \sqrt{h} h^{ab} \rightarrow \sqrt{h} h^{ab} \right\} \text{ invariant}$$

Symmetries:

- Reparametrization invariance (on the worldsheet: $\sigma^i \rightarrow \sigma'^i(\sigma^i)$)
- Weyl ~~symmetry~~ invariance

try to eliminate h_{ab} : $\frac{\delta S}{\delta h_{ab}} = 0 \Rightarrow H_{ab} - \frac{1}{2} h_{ab} (h^{cd} H_{cd}) = 0$

$$\text{so: } h_{ab} = \frac{2}{h(H_{ab})} H_{ab}, \quad |h| = \frac{4}{(tr H_{ab})^2} |H|$$

$$S = \frac{T}{2} \int \sqrt{h} tr H_{ab} = \frac{T}{2} \int \frac{2 \sqrt{H}}{(tr H)} tr H = \frac{T}{2} \int \sqrt{H} d^2 \sigma =$$

$$= T \int \sqrt{H} d^2 \sigma, \text{ Nambu - Goto action.}$$

H_{ab} is the induced metric:

- the restriction of $g_{\mu\nu}$ to the worldsheet



$$g_{\mu\nu} V^\mu W^\nu = g_{\mu\nu} V^a \partial_a X^\mu W^b \partial_b X^\nu = (g_{\mu\nu} \partial_a X^\mu \partial_b X^\nu) V^a W^b = H_{ab} V^a W^b$$

Like in GR:

$$\frac{\delta}{\delta g_{\mu\nu}} \int d^4 x \sqrt{-g} g^{\mu\nu} R_{\mu\nu} = (R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R)$$