

# Supersymmetry

and other theories of  
Dark Matter Candidates

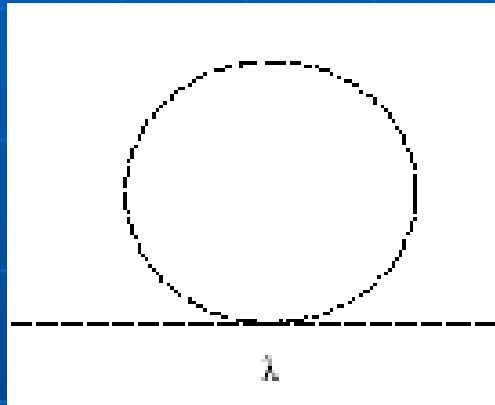
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798G Presentation  
3/1/07

# Overview

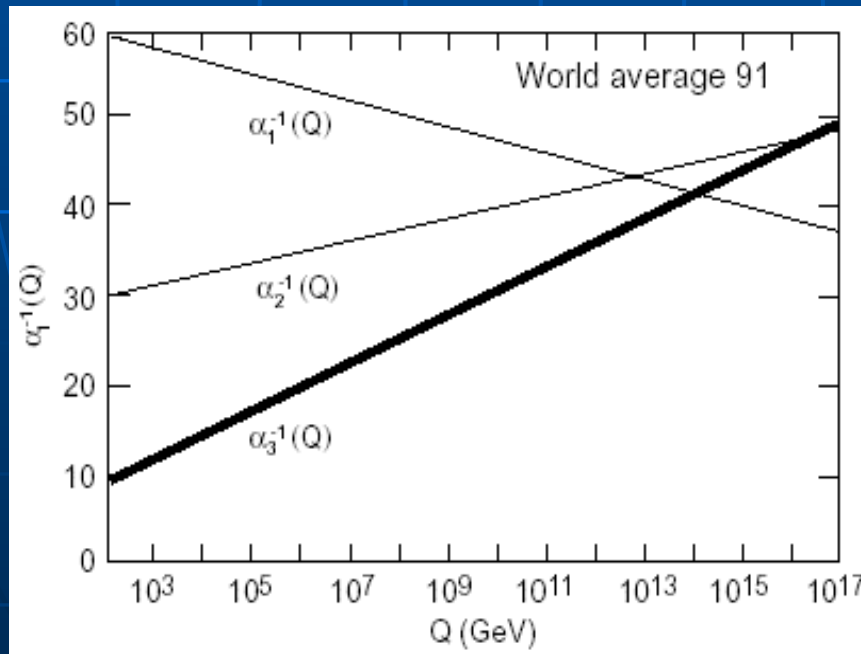
- Why bother with a new theory?
- Why is Supersymmetry a good solution?
- Basics of Supersymmetry
- Why this leads to Dark Matter Candidates
- Other theoretical basis for Dark Matter Candidates



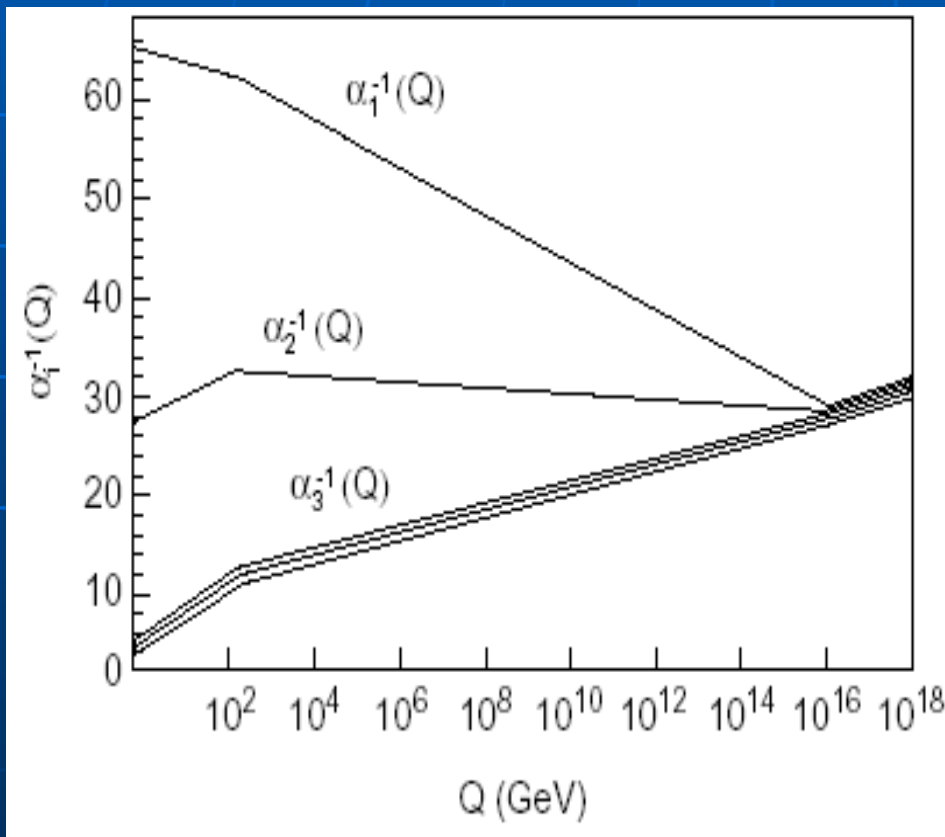
# Hierarchy Problem



- Quadratic divergence in Higgs loops
- Can't renormalize Higgs mass without significant cancelations
- Electroweak scale vs. Planck scale



# How does SUSY Help?



- Unification of Forces
  - Energy dependent coupling parameters meet with SUSY
- Higgs Mass
  - Divergences cancel due to opposite sign contributions
- Abundance of matter over antimatter, size and age of universe, proton decay, ...

# Basics of SuSy

- What is a symmetry?
  - Lagrangian remains unchanged under a transformation of the fields
  - Ex: SM Lagrangian doesn't change if fields are "boosted" (transformed to moving frame using Spec. Rel.)
- What is a SUPER symmetry?
  - Transformation of fields include fields of opposite nature
  - Terms cancel in Lagrangian when all fields transformed, leaving it changed by only a derivative

# The Mathematics

## ■ Wess-Zumino Model

$$\mathcal{L} = \frac{1}{2}(\partial_\mu A)^2 + \frac{1}{2}(\partial_\mu B)^2 + \frac{i}{2}\bar{\psi}\not{\partial}\psi - \frac{1}{2}m\bar{\psi}\psi - \frac{1}{2}m^2A^2 - \frac{1}{2}m^2B^2 \\ + mgA(A^2 + B^2) - \frac{1}{2}g^2(A^2 + B^2)^2 - ig\bar{\psi}\overset{\textcircled{1}}{A}\psi + ig\bar{\psi}\overset{\textcircled{2}}{\gamma_5}B\psi$$

$$A \rightarrow A' = A + \delta A = A + \bar{\alpha}QA$$

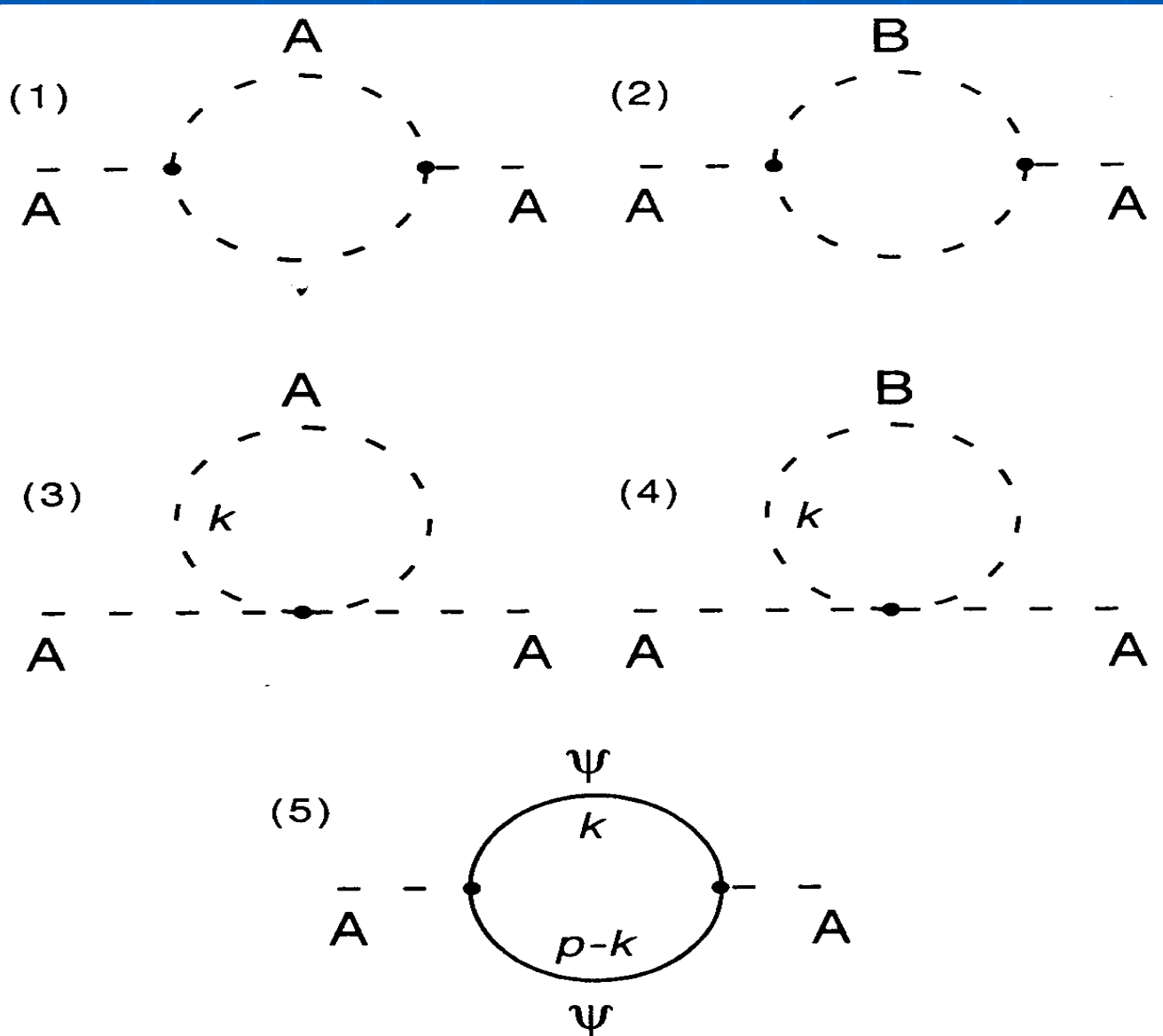
$$\delta A = i\bar{\alpha}\gamma^5\psi$$

$$\delta B = -\bar{\alpha}\psi$$

$$\delta\psi = F\alpha - iG\gamma^5\alpha + (\not{\partial}\gamma^5A)\alpha + i(\not{\partial}B)\alpha$$

where  $F = mA - g(A^2 - B^2)$  and  $G = mB - 2gAB$ .

# Eliminating Quadratic Divergences





# Naming Scheme

- Fermions get an "s" in front
- Bosons get an "-ino" at the end
- Symbols get a "~" on top.

<i>Particle</i>	<i>Name</i>	<i>Feels These Forces<sup>a</sup></i>	<i>Mediates These Forces<sup>b</sup></i>	<i>Superpartner</i>
$e, \mu, \tau$	charged leptons (electron, muon, tau)	EM, W	—	sleptons $\tilde{e}, \tilde{\mu}, \tilde{\tau}$ (selectron, smuon, stau)
$\nu_e, \nu_\mu, \nu_\tau$	neutrinos	W	—	sneutrinos $\tilde{\nu}_e, \tilde{\nu}_\mu, \tilde{\nu}_\tau$
$u, c, t$	up, charm, top quarks	EM, W, S	—	squarks $\tilde{u}, \tilde{c}, \tilde{t}$
$d, s, b$	down, strange, bottom quarks	EM, W, S	—	squarks $\tilde{d}, \tilde{s}, \tilde{b}$
$\gamma$	photon	$\epsilon$	EM	photino <sup>d</sup> $\tilde{\gamma}$
$W^\pm$	weak boson	EM, W	W	Wino <sup>d</sup> $\tilde{W}^\pm$
$Z$	weak boson	W	W	Zino <sup>d</sup> $\tilde{Z}$
$g$	gluon	S	S	gluino $\tilde{g}$
$G$	graviton	GR	GR	gravitino $\tilde{G}$
$h$	Higgs boson <sup>e</sup>	W	generates mass	higgsino <sup>e</sup> $\tilde{h}$

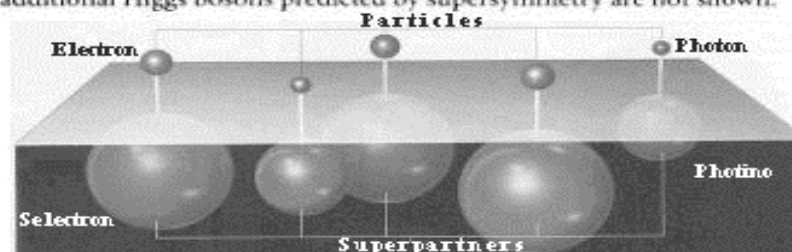
<sup>a</sup>All particles feel the gravitational force.

<sup>b</sup>EM = electromagnetic force, W = weak force, S = strong force, GR = gravitational force.

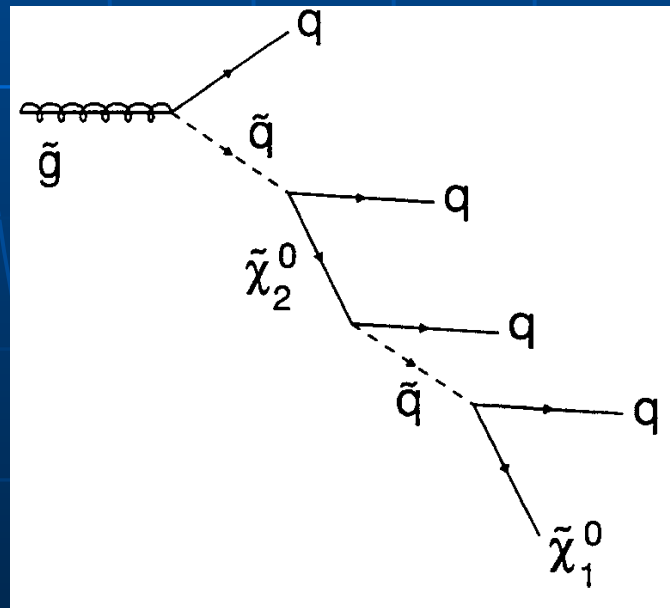
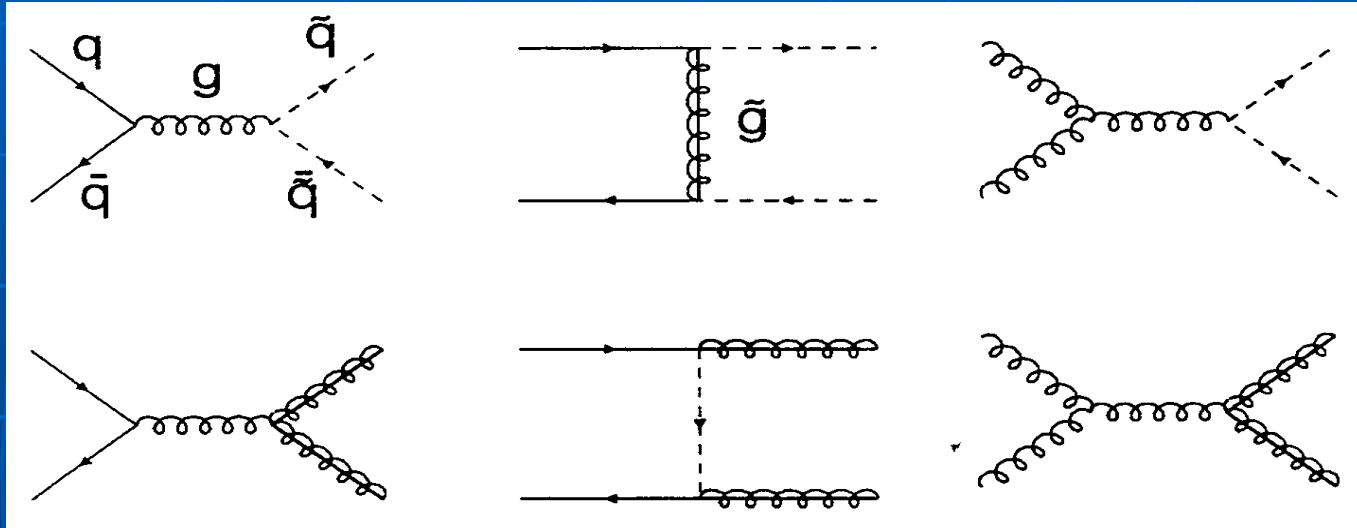
<sup>c</sup>Photons feel only the gravitational force, but they interact with all electrically charged particles.

<sup>d</sup>Mixtures of these particles form charginos and neutralinos (Appendix C).

<sup>e</sup>The additional Higgs bosons predicted by supersymmetry are not shown.

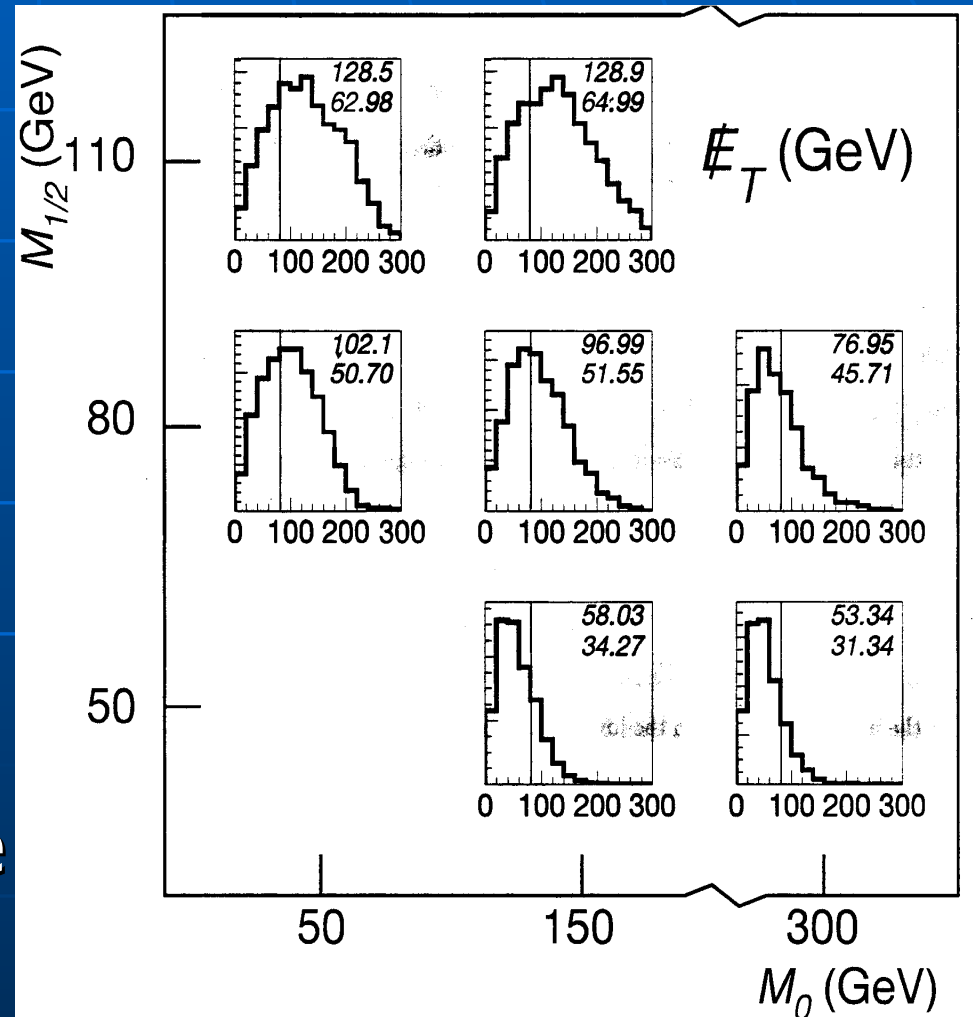


# A typical SuSy decay mode



# How could we see it?

- The Lightest Supersymmetric Particle (LSP) neutral, so only interact via weak force
- Very hard to detect, so seen as Missing Transverse Energy in decays



# Common Types of SuSy

## ■ MSSM

- Minimal Supersymmetric Model
- Adds fewest number of new particles
- Adds new conserved quantum number:
  - $R = (-1)^{3(B-L)+2S}$

## ■ mSUGRA

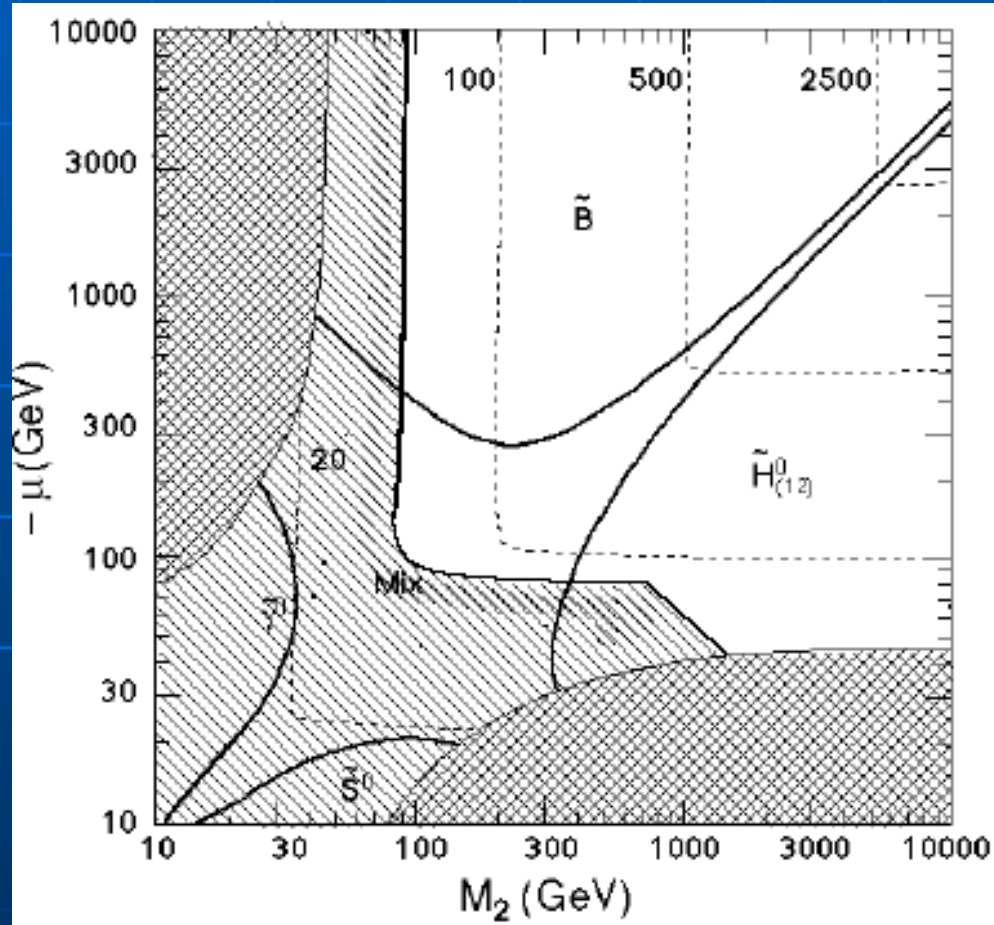
- Minimal Supergravity
- Decreases # of parameters from  $>100$  to 5 by grouping some together and requiring they be equal

# Why is this a Dark Matter Candidate?

- Dark matter must interact only via the weak force
  - If electrically or color charged would see it in Nuclear effects such as anomalously heavy ions, "Very hard to hide Baryons."
- If R quantum # conserved the Lightest Supersymmetric Particle (LSP) would be stable
- LSP must be relatively massive since hasn't been detected yet

# What are the Candidates?

- Gravitino
  - Very difficult to confirm or exclude
- Sneutrino
  - In MSSM has been excluded by direct and indirect searches
- Neutralino
  - Could be photino, bino, or Higgsinos



$$\chi = \alpha \tilde{B} + \beta \tilde{W}^3 + \gamma \tilde{H}_1 + \delta \tilde{H}_2$$

# What if we don't see SUSY?

- There must be modifications to SM
- SUSY provides an EXTENSION of SM, leaving most of it intact
- Time to scrap SM and start over?
- New theories that avoid quadratic divergences (renormalization?), explain hierarchy problem, unify forces



# Other theories that provide DMC

## ■ WIMPs

- Weakly Interacting Massive Particles
- Generic name for DMC that interacts only via weak force
- Could have not seen it yet due to lifetime, mass, small production cross section

## ■ Axions

- Strong Lagrangian not unchanged under CP transformation, yet no CP violation seen
- Solution involves introduction of mechanism similar to Higgs, but with a MASSIVE Goldstone boson – Axion.



# Summary

- Don't forget, there is very little – if any – evidence for SUSY so far.
- Provides solutions to many problems of the SM, but requires many arbitrary parameters
- Gives a clear prediction for mass ranges of spartners,  $(m_B^2 - m_F^2) < 1$  TeV, so we'll know in a few years ...

# References

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